## **Explanatory Supplement To The Astronomical Almanac Pdf**



- contents- chapter 1: introduction- chapter 2: physical principles- chapter 3: celestial mechanics- chapter 4: positional astronomy- chapter 5: ephemerides- chapter 6: the fundamental celestial reference systemchapter 7: astrometry and positional astronomy- chapter 8: time- chapter 9: terrestrial coordinates and the rotation of the earth- chapter 10: positions- chapter 11: orbital ephemerides of the sun, moon, and planetschapter 12: planetary satellites and rings- chapter 13: physical ephemerides- chapter 14: eclipses of the sun and moon- chapter 15: astronomical phenomena- chapter 16: stars and stellar systems- chapter 17: computational techniques- chapter 18: calendars- chapter 19: bases and constants- chapter 20: glossary for a more detailed description of the principles of spherical trigonometry, see the excellently written article by m. radice: "cartesian-spherical coordinate systems" (2009) in the journal of astronomical history and heritage, 12(2), pp. 29-54. it is available as a web page, the section "theoretical basis of spherical trigonometry" in that paper gives a concise summary of the theory. the astronomical coordinates of the sun are the azimuth and the elevation of the sun. measured from a fixed point in the sky (the reference direction). these are the angles of the right ascension and the declination of the sun. the right ascension is the angle between the plane of the ecliptic and the line from the sun to the observer, and the declination is the angle between the normal to the plane of the ecliptic and the projection of the line from the sun to the observer on the plane of the ecliptic. the celestial equator is the ecliptic, the azimuth, the elevation, and the right ascension are measured in the same direction and are recorded together. for most applications, only the azimuth and the elevation are needed. the right ascension of the sun is also known as solar time. the time of the sun is the fraction of the sidereal day to which the sun moves in the course of a single sidereal day, the sun's time does not change with time, and is zero at the moment of the vernal equinox, the time of the sun is also called the sun's mean motion (or mean motion of the sun) and denoted by e or ε.

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the first chapter of the explanatory supplement describes the celestial reference frame and how celestial coordinates are specified. the second chapter describes the terrestrial reference frame and how terrestrial coordinates are specified, the third chapter is devoted to the motion of the earth with respect to the barycenter of the solar system. (the use of the term "barycenter" is inexact, since the reference frame is centered on the earth.) the fourth chapter describes the motion of the earth with respect to the center of the visible universe. the fifth chapter describes celestial coordinates and celestial reference systems in detail, the precise values of constants in the iau system are given in the explanatory supplement to the astronomical almanac, the values of astronomical constants in the iau system are computed by statistical methods and therefore have higher accuracy than the older fixed-value tabulations. in particular, the values of astronomical constants differ by more than 0.2% from the values of the tabulations of twinsun, the changing of the constants is described by a set of constants, the e-terms of aberration (eg. (13) of seidelmann, 2006), the absolute values of the fundamental constants of the iau system and the e-terms of aberration are given in the explanatory supplement to the astronomical almanac. the e-terms of aberration are used to determine corrections to the observational time scales, which are used in the determination of the astronomical constants. following the si system of units, all measured quantities are expressed in terms of the basic unit of length, the meter, the meter is also the basic unit of distance, the length of a celestial meridian; it is usually defined as the length of the path of light in vacuum from the center of the earth to a point where its speed is the same as that of light, the astronomical unit (au) is the mean length of the meridians of the earth and the celestial sphere, this is the distance from the center of the earth to the center of the sun, taking into account the earth's motion and the curvature of the celestial sphere, the astronomical unit (au) is not the same as the astronomical unit (1 au is 149.598,262,730 meters), 5ec8ef588b

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